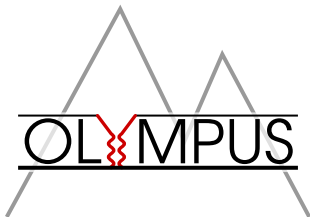


Status of the OLYMPUS Analysis

D.K. Hasell

M.I.T.

PRC Review



Outline

- 1 Status of Analysis
 - Optical Survey
 - Magnetic Field Mapping
 - TOF Detectors
 - 12° Luminosity Monitors
 - Symmetric Møller / Bhabha Detector
 - Luminosity
 - Monte Carlo
 - Track Reconstruction
 - Analysis of February Data
- 2 Additional Activities
 - Disassembly and Shipping
 - Moving Computers and Reprocessing
 - Collaboration Meetings and Workshops
 - NIM Paper on OLYMPUS Experiment
- 3 Summary and Timeline
 - Summary
 - Timeline

Optical Survey

Measurements taken by [DESY survey group](#)

- raw data \rightarrow 2 angles and a distance relative to laser tracker
- targets on detectors, frames, walls, floor, the DORIS ring, ...
- offsets for target mount, target prism, angle relative to laser tracker

Different concepts or goals for the survey

Raw data analysed for OLYMPUS - [Jan Bernauer \(MIT\)](#)

- combined many sets of raw data, fixed inconsistencies
- performed global fit
- determined detector locations
- accuracy $< 100 \mu\text{m}$ for most points



GDML

Geometry Description Markup Language - [Colton O'Connor \(MIT\)](#)

- detailed description of detectors and infrastructure
 - target, scattering chamber, beamline, and shielding - fully implemented
 - toroid coils - to within 2 cm (placed according to field map fit)
 - wire chambers, TOF, GEM, SYMB - fully implemented
 - MWPC - active area fully implemented, electronics nearly finished
 - Lead Glass - active area fully implemented, need support structure
 - SiPMs - to within 1 cm along 12° line
- most elements based on optical survey results
- used for both analysis and Monte Carlo



Magnetic Field Mapping

March - April, 2013 measured magnetic field at 36,000 points

- group effort [DESY](#), [HU](#), [MIT](#), [PNPI](#), [Yerevan](#)
- two 6 hour shift per day, 3 persons per shift
- primary XYZ table (± 100 mm X, ± 100 mm Y, $\pm 3,000$ mm Z)
- secondary XY table (± 500 mm X, ± 500 mm Y)
- laser tracker only available for initial setup in left and right sectors
- several other setups required

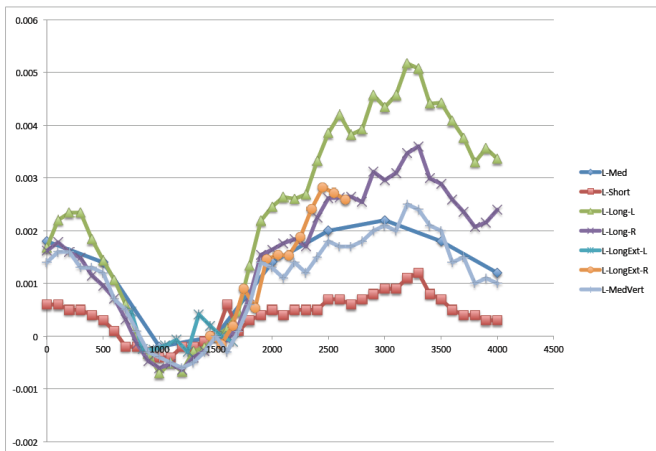
Probe position oscillated $\pm 1-2$ mm with Z

- measure oscillation for each setup with theodolite total station
- for each scan measure start and end points
- correct for oscillation assuming pattern fixed for each setup

Analysis by [Axel Schmidt](#), [Brian Henderson](#), [Colton O'Connor](#) (MIT)



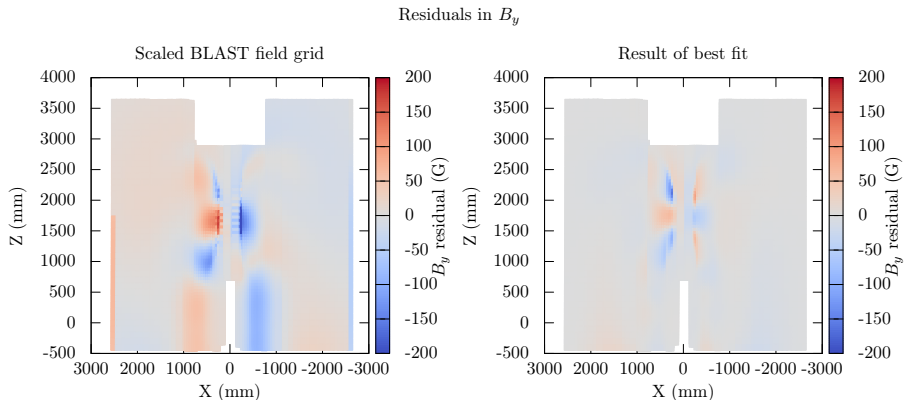
Oscillation of 3D Hall Probe



Depended on left/right sector, long/medium/short rod, extension



Residuals in Magnetic Field



Global fit to magnetic field data using Biot-Savart calculation

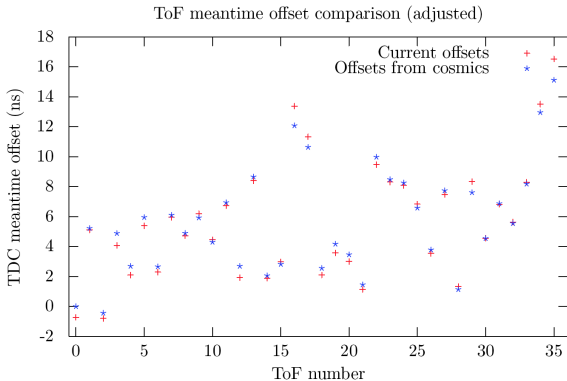
- average residual at a point **< 19 G**
- generate regular grid, 420,000 points, for analysis and Monte Carlo



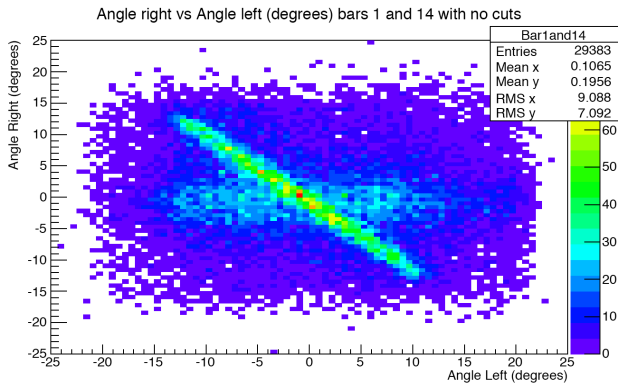
Time of Flight Detectors

Calibration and analysis by [ASU](#), [MIT](#), and [Yerevan](#)

- time offsets, ADC pedestals, gains, attenuation lengths



TOF Coplanarity - Lauren Ice (ASU)



Azimuthal angle from top/bottom PMT time difference

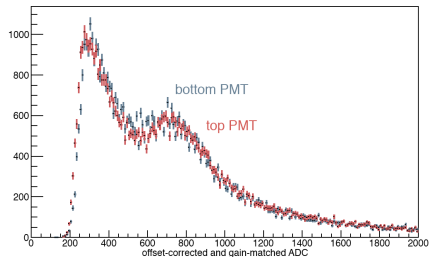
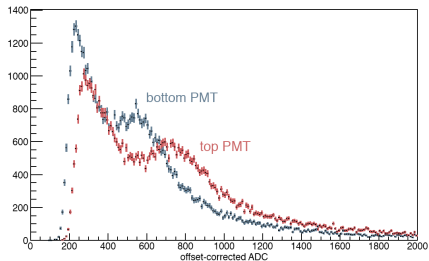
- clear correlation above background and random coincidences



ADC Gain Analysis - Rebecca Russell (MIT)

$$ADC_{i,t} = G_i f E_{dep} e^{\frac{(y-0.5L_i)}{l_i}} + C_{i,t}$$

$$ADC_{i,b} = \alpha_i G_i (1-f) E_{dep} e^{\frac{-(y+0.5L_i)}{l_i}} + C_{i,b}$$



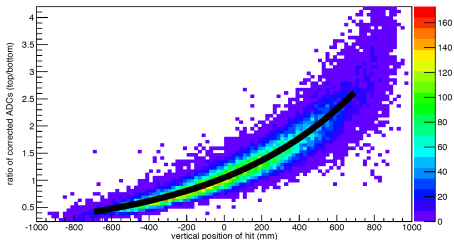
- α_i relative gain matching of top / bottom PMTs on each TOF bar
- $f \sim 0.5$ mostly independent of track angle and vertical position



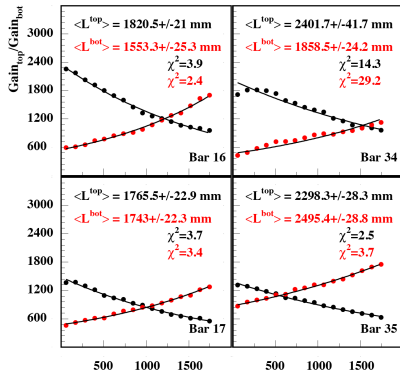
TOF Attenuation Lengths

Two approaches **Rebecca Russell (MIT)** and **Norik Akopov (Yerevan)**

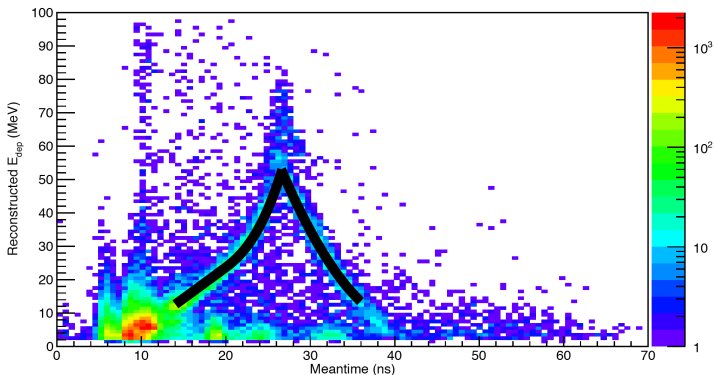
- comparable results



$$\frac{\alpha_j (ADC_{i,t} - C_{i,t})}{ADC_{i,b} - C_{i,b}} = \left(\frac{f}{1-f} \right) e^{2y/l_i}$$



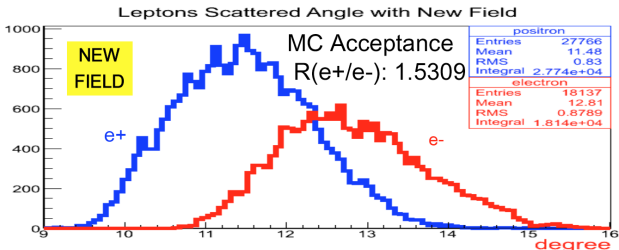
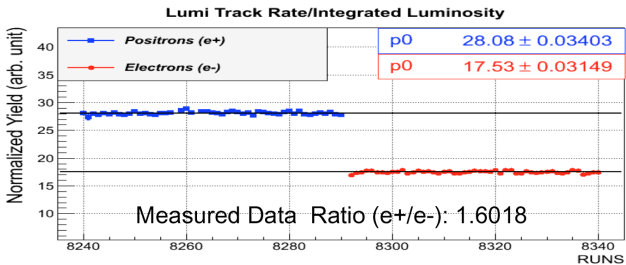
Absolute TOF Gain from Protons - Rebecca Russell (MIT)



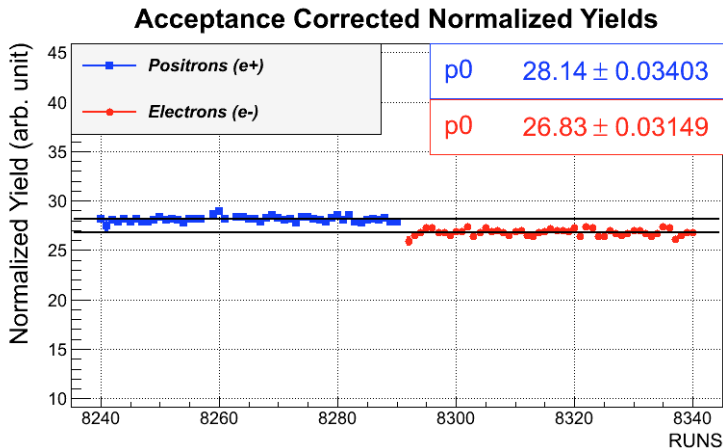
- correction for ADC overflow
- poor statistics for forward bars
- comparable results using muons from cosmic runs



GEM Detector Rates - Özgür Ates (HU)



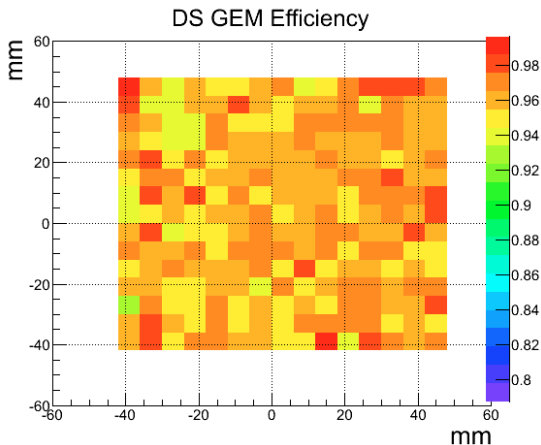
GEM Acceptance Corrected Detector Rates



Ratio 1.049 ± 0.005



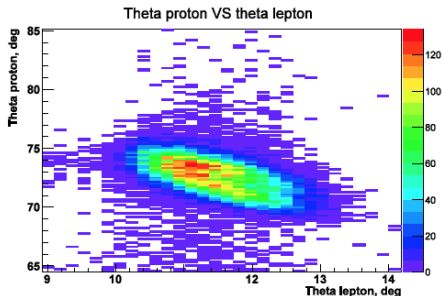
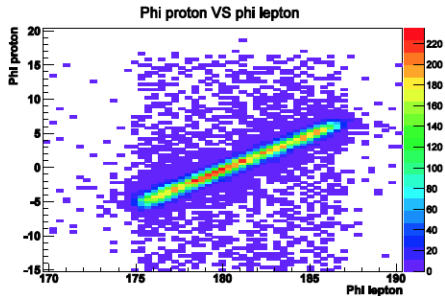
GEM Detector Efficiency - Özgür Ates (HU)



- efficiency $\sim 95\%$ in all detectors
- good agreement between MC simulation and measured rates

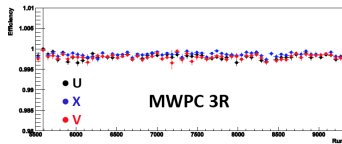
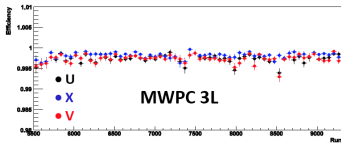
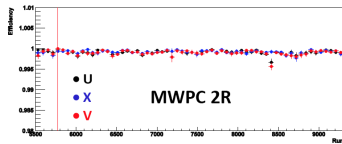
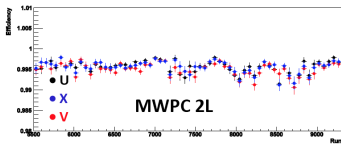
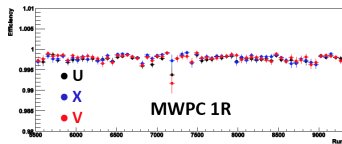
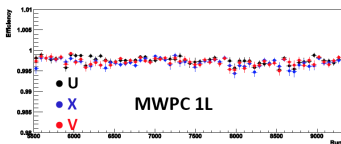


MWPC - Denis Veretennikov (PNPI)



- reasonable ep elastic resolution
- ratio of e^+p/e^-p rates agree with MC to $\sim 5\%$

MWPC Efficiency - Denis Veretennikov (PNPI)



● efficiency $\sim 99\%$



Symmetric Møller / Bhabha Detector

Analysis and Monte Carlo - [Roberto Pérez Benito \(Mainz\)](#)

- using optical survey data to account for geometric effects
- correcting for beam position and slope
- Monte Carlo simulation at tree level working
 - higher order effects being discussed with theory group at Mainz
- dynamic non-linearities in ADC being studied

Good results shown in April seem to have changed

- naive approach gave good agreement
- corrections, particularly geometric effects, cause discrepancy
- needs to be investigated further



Symmetric Møller / Bhabha Detector

	Geometrical Cross section [mbarn]	Loss cross section %	Coincidence cross section [mbarn]
Møller (MC run 9505)	0.00538	23.2	0.00414
Bhabha+Annihilation (MC run 9429)	0.00322	29.3	0.00227
ratio	1.67		1.82

	Measured cross section [mbarn]
Møller (run 9505)	0.00471
Bhabha + Annihilation (run 9429)	0.00294
ratio	1.60

Run 9505 Møller

dtc luminosity : $7.0667e+35 \text{ cm}^{-2}$

møller events 3326573

measured cross section 0.00471 mbarn

Run 9429 Bhabha+Annihilation

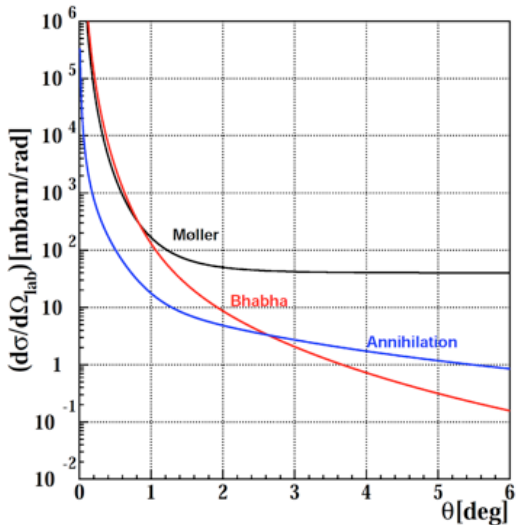
dtc luminosity : $1.3511e+35 \text{ cm}^{-2}$

bhabha+annihilation events 396772

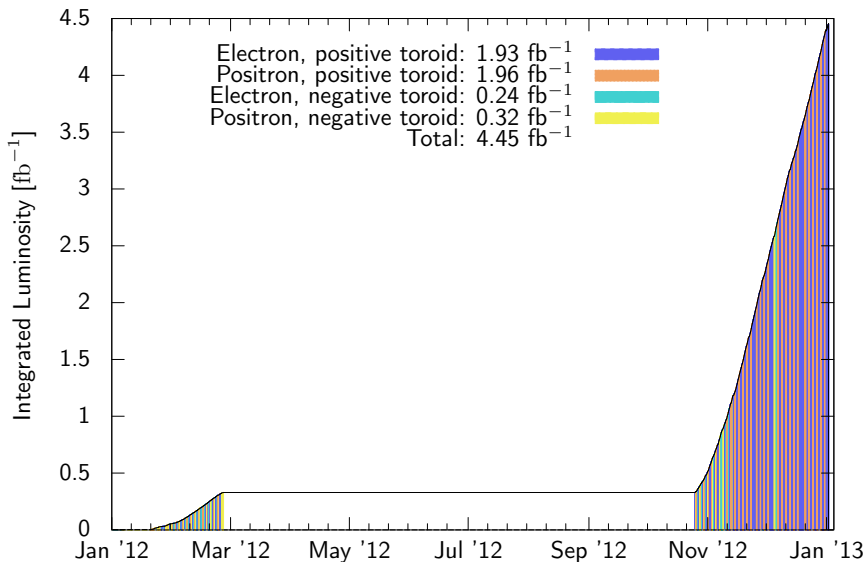
measured cross section 0.00294 mbarn



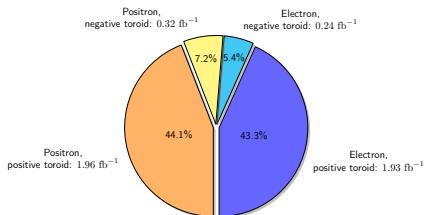
Symmetric Møller / Bhabha Detector



Luminosity from Slow Control - Jan Bernauer (MIT)



Luminosity



Most data with positive magnetic field

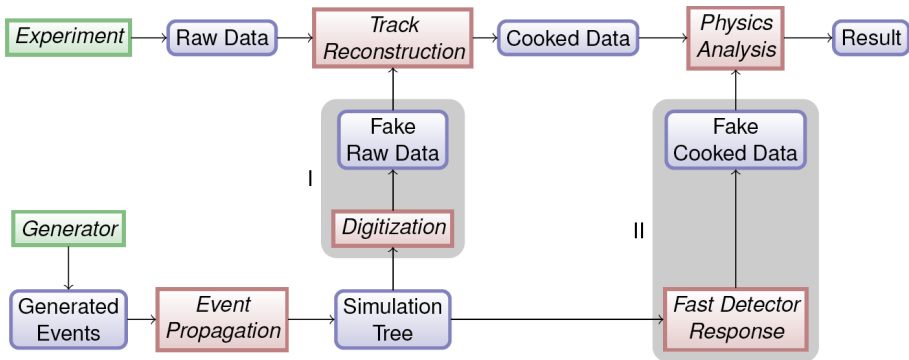
- negative polarity bends electrons into the wire chambers
- wire chambers unable to handle noise rate

Luminosity work in progress - [Jürgen Diefenbach \(Mainz\)](#)

- comparing 12° luminosity detector and Møller/Bhabha results
- started using new MC for 12° for acceptance studies



Monte Carlo



- new framework for Monte Carlo integrated in Cooker framework
- all groups expected to contribute to analysis and MC repository
 - primarily [MIT](#), [ASU](#)
 - some from [HU](#), [Mainz](#), [PNPI](#)

Monte Carlo Generators and Digitization

Generators - MIT, ASU, DESY

- basic generators and pion generator - 100 %
- implementing radiative effects in generators
 - internal bremsstrahlung generator - 90 %
 - port of Novosibirsk generator almost - 90 %
 - common effort with JLAB and Novosibirsk
 - DESY version of Maximon and Tjon done but not in repository

Digitization

- digitization of all components underway
 - TOF 90 %
 - WC 50 %
 - GEM 70%
 - MWPC 10 %
 - SYMB
 - trigger 100 %



Track Reconstruction - Jan Bernauer (MIT)

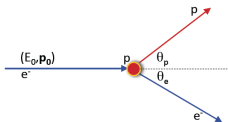
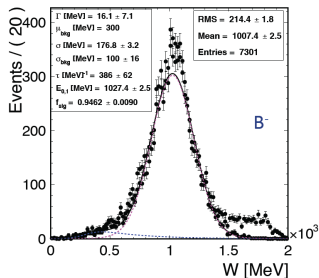
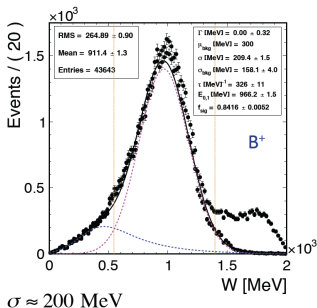
- will use Elastic-Arms-Algorithm
 - global fit version of Kalman filter
 - also have Kalman filter
 - early tests encouraging
 - decided to wait until optical survey, magnetic field map, and data reprocessing completed and implemented
 - need to find time-to-distance for each wire from data
 - must also tune algorithm (cut-offs, etc.)
 - replaced path swimming with spline lookup to improve speed
- ready to begin



Analysis of February Data - Nicola D'Ascenzo (DESY)

Using a small subset of February runs with tracking results

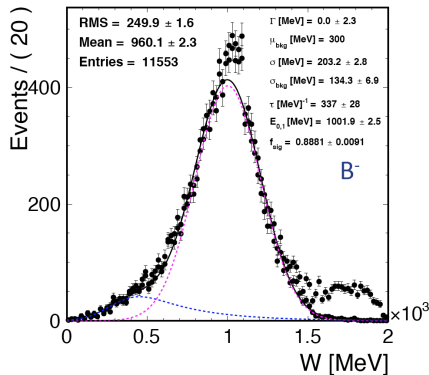
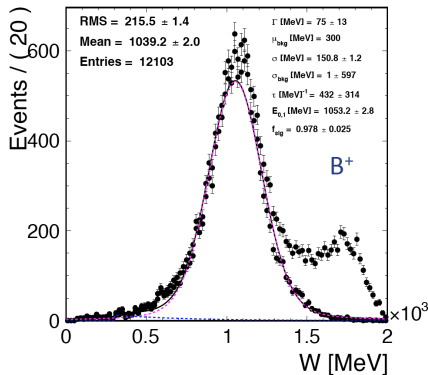
Reconstruction from angles and energy



$$W = \sqrt{m_p^2 + 2m_p(E_0 - E_e) - 4E_0E_e \sin^2 \frac{\theta_e}{2}}$$

For elastic events $W = m_p$

Analysis of February Data - Nicola D'Ascenzo (DESY)



- *ep* elastic scattering events clearly seen
- based on old, first attempt at track reconstruction
- missing tracks from B^- runs



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Disassembly and Shipping

- after magnetic field mapping and optical survey complete
- started disassembly of OLYMPUS experiment
- useful items (electrons, power supplies, vacuum equipment, computers, etc.) recovered by various institutes
- mechanical structure of the experiment and the magnet coils scrapped



Moving Computers and Reprocessing - Bonn

- after cosmic data run computers moved to Bonn
- temporarily without access to nominal run database, wiki, data, etc.
- converting raw ZEBRA data to ROOT files used for analysis
 - some channels mis-mapped
 - first reprocessing had some problems
 - solved now and most data converted
- data will be available at MIT and DESY



Collaboration Meeting at MIT July 8–11, 2013

- very successful and productive meeting
- 2 mornings of normal meeting - updates and status reports
- rest of meeting in workshops
 - concentrated on specific areas
 - tutorial on using the Cooker framework
 - tutorial on using the new Monte Carlo
 - special sessions on MC, Tracking, TOF, GEM, MWPC, etc.
 - reviewed analysis, discussed work to do, simulation, digitization, . . .



Next Collaboration Meeting at DESY, October 17-18, 2013

- ahead of PRC meeting October 24-25, 2013
- working on radiative correction workshop at Orsay, October 7-8, 2013



NIM Paper on OLYMPUS Experiment

- description of experimental configuration and operation
 - DORIS accelerator
 - internal, gas target and vacuum system
 - detector system
 - trigger, data acquisition, slow control, data quality control
 - operation of experiment
- before experts disappear or forget
- available for citation
- paper fully written
- comments, suggestions, details until September
- final version ready for approval by October



Summary

- a lot of work has been done
- field mapping, optical survey, GDML, and data reprocessing all done
- analysis well advanced on TOF, GEM, and MWPC ahead of tracking
- track reconstruction is set to begin
- Monte Carlo incorporated into Cooker framework
- several components digitized
- first radiative corrections added to generators, planning workshop
- luminosity determinations have to be improved/understood
- main effort is at MIT (ASU student was at MIT)
- other efforts isolated and/or need to be integrated better
- NIM paper ready for submission in October



Timeline

- first passes on track reconstruction in September
- initial tracking results in October
- detailed analysis by April, 2014
- final results by end 2014

